

Event-by-event fluctuations in hydrodynamical description of heavy-ion collisions*

C.E. Aguiar^a[UFRJ] Instituto de Física/UFRJ, C.P. 68528, 21945-970 Rio de Janeiro - RJ, Brazil, Y. Hama^b[USP] Instituto de Física/USP, C.P. 66318, 05389-970 São Paulo - SP, Brazil, T. Kodama^c[UFRJ] and T. Osada^d[USP]

^a[

^b[

Effects caused by the event-by-event fluctuation of the initial conditions in hydrodynamical description of high-energy heavy-ion collisions are investigated. Non-negligible effects appear for several observable quantities, even for a fixed impact parameter \bar{b} . They are sensitive to the equation of state, being the dispersions of the observable quantities in general smaller when the QGP phase appears at the beginning of hydrodynamic evolution than when the fluid remains hadron gas during whole the evolution.

1. INTRODUCTION

In usual hydrodynamic description of high-energy heavy-ion collisions, one customarily assumes some highly symmetric and smooth initial conditions, which correspond to mean distributions of velocity, temperature, energy density, etc., averaged over several events. However, our systems are not large enough, so large fluctuations are expected. What are the effects of the event-by-event fluctuation of the initial conditions? Are they sizable? Do they depend on the equation of state? Which are the most sensitive variables? These are some questions which arise regarding such an initial-state fluctuation, and we try to shed some light on these matters in the present study[1].

2. METHOD OF STUDY

In order to study the problem stated above, first we generate events by using the NeXus event generator[2], from which initial conditions are computed at the time $\tau = 1$ fm. Then, the hydrodynamic equations are solved, starting from these initial conditions, assuming some equation of state (EoS). To see the EoS dependence of the effects we are treating, we consider two different EoS's[3]:

1. Resonance Gas (RG): $c_s^2 = 0.2$;

*Work supported in part by FAPESP (contract nos. 2000/04422-7 and 98/00317-2), FAPERJ (contract no.E-26/150.942/99), PRONEX (contract no. 41.96.0886.00) and CNPq-Brasil.

$$2. \text{ QGP+RG: } c_s^2 = \begin{cases} 0.2, & \varepsilon < 0.28 \text{ GeV/fm}^3, \\ 0.056/\varepsilon, & \text{mixed phase}, \\ 1/3 - 4B/3\varepsilon, & \varepsilon > 1.45 \text{ GeV/fm}^3. \end{cases}$$

The resolution of the hydrodynamic equations deserves some special care, since our initial conditions do not have any symmetry nor they are smooth. We adopt the so-called smoothed-particle hydrodynamic (SPH) approach[4], first used in astrophysics and which we have previously adapted for heavy-ion collisions[5], a method flexible enough, giving a desired precision. The main characteristic of SPH is the parametrization of the flow in terms of discrete Lagrangian coordinates attached to small volumes (called “particles”) with some conserved quantity. In the present work, besides the energy and momentum, we took the entropy as our conserved quantity. Then, its density (in the space-fixed frame) is parametrized as

$$s^*(\mathbf{x}, t) = \sum_i^N \nu_i W(\mathbf{x} - \mathbf{x}_i(t); h), \quad (1)$$

where

$$\begin{cases} W(\mathbf{x} - \mathbf{x}_i(t); h) \text{ is a normalized kernel;} \\ \mathbf{x}_i(t) \text{ is the } i\text{-th particle position, so the velocity is } \mathbf{v}_i = d\mathbf{x}_i/dt; \\ h \text{ is the smoothing scale parameter;} \end{cases}$$

and we have

$$S = \int d^3\mathbf{x} s^*(\mathbf{x}, t) = \sum_i^N \nu_i. \quad (2)$$

The equations of motion are then written as the coupled equations

$$\frac{d}{dt} \left(\nu_i \frac{P_i + \varepsilon_i}{s_i} \gamma_i \mathbf{v}_i \right) + \sum_j \nu_j \left[\frac{P_i}{s_i^{*2}} + \frac{P_j}{s_j^{*2}} \right] \nabla_i W(\mathbf{x}_i - \mathbf{x}_j; h) = 0. \quad (3)$$

Following this procedure, we computed some observable quantities, event-by-event, for $\sqrt{s}=130 \text{ AGeV } Au + Au$ collisions. The results are presented in the next Section.

3. RESULTS

3.1. Elliptic flow coefficient v_2

Having solved the coupled equations (3), we have computed the particle spectra at $T = m_\pi$ and from which the elliptic flow coefficient v_2 on an event-by-event basis. In Figure 1, we show its distributions for a fixed impact parameter b , for the two EoS considered. As expected, v_2 exhibits a large fluctuation, which depends on the EoS. One should take care in looking at this Figure that our b is the true impact parameter (not determined in the way experimentalists do), so for instance in the RG case, there are some events with negative v_2 , which experimentally would not appear. As for the average values $\langle v_2 \rangle$, it is almost independent of the EoS. This is shown in Figure 2, where $\langle v_2 \rangle \pm \delta v_2$ is plotted as function of the centrality and compared with data[6]. It is seen that $\langle v_2 \rangle$ reproduces well the experimental trend, whereas the dispersions δv_2 are much wider than the experimental errors. As for the EoS dependence, δv_2 is smaller when QGP is produced.

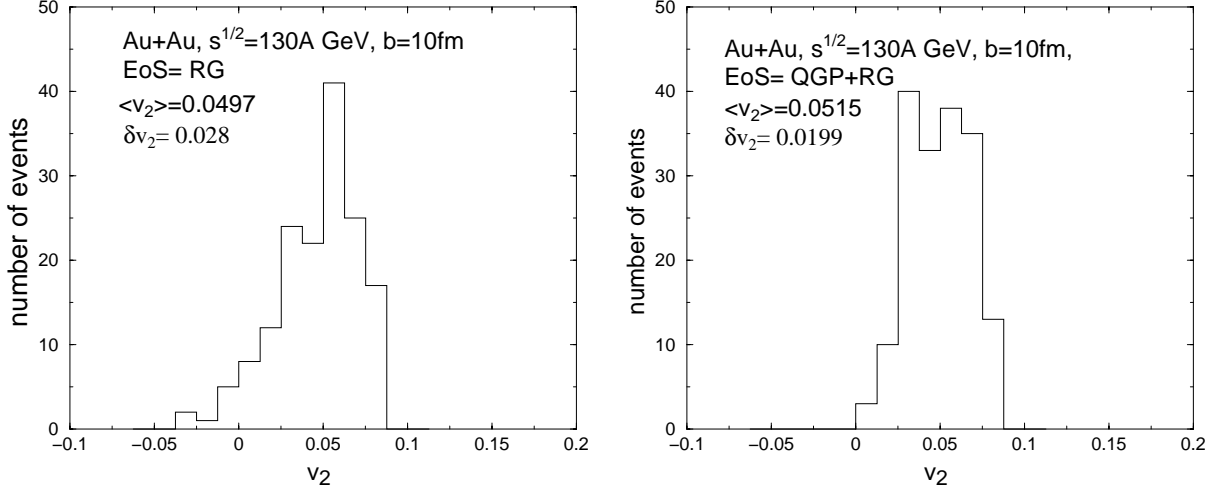


Figure 1. Distribution of elliptic-flow coefficients v_2 at $b = 10$ fm for two EoS.

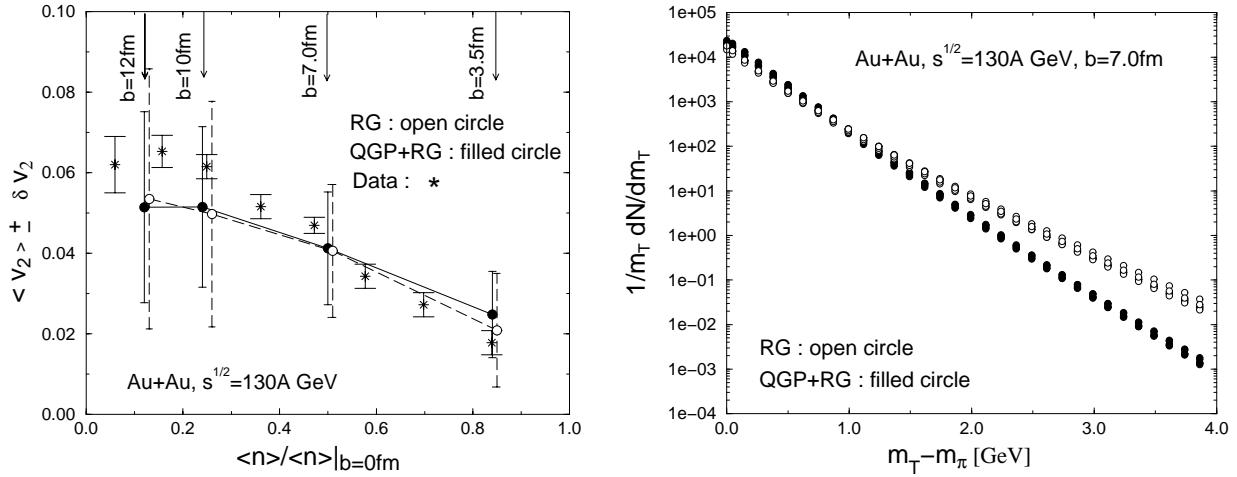


Figure 2. EoS dependence of $\langle v_2 \rangle \pm \delta v_2$ as function of the centrality, compared with ($b = 7.0$ fm). data[6].

3.2. m_T distributions

In Figure 3, we show the m_T distributions for 5 events. As expected, m_T distributions are in general steeper when QGP is produced. As for fluctuations, the resultant fluctuation in m_T spectrum (or in the slope parameter $\delta \tilde{T}$) is very small.

3.3. Multiplicity fluctuation in the mid-rapidity region

Table 1 summarizes the results of our study on multiplicity fluctuation in the mid-rapidity region. It is seen that *i*) as $b \rightarrow 0$, $\langle n_\pi \rangle$ becomes much larger with the QGP EoS; *ii*) δn_π shows the same tendency in this limit; *iii*) As for the ratio $\delta n_\pi / \langle n_\pi \rangle$, it is not sensitive to the EoS.

Table 1

EoS dependence of the multiplicity fluctuation in two different rapidity intervals $-\Delta y < y < +\Delta y$, as function of the impact parameter b .

b [fm]	EoS	# of events	$\Delta y=1.875$			$\Delta y=3.00$		
			$\langle n \rangle$	δn	$\delta n / \langle n \rangle$	$\langle n \rangle$	δn	$\delta n / \langle n \rangle$
3.5	RG	44	1029.7	46.2	0.045	1623.2	68.7	0.042
	QGP	38	1553.0	80.9	0.052	2544.6	129.0	0.051
7.0	RG	55	613.3	49.5	0.081	977.4	71.6	0.073
	QGP	58	926.1	81.1	0.087	1530.5	123.7	0.081
10.0	RG	166	312.8	43.0	0.137	506.1	65.8	0.130
	QGP	180	437.5	66.5	0.151	740.7	103.9	0.140
12.0	RG	79	162.8	35.6	0.219	268.9	56.2	0.209
	QGP	100	220.1	52.8	0.240	379.8	85.2	0.224

4. CONCLUSIONS AND OUTLOOK

The present study shows that the effects of the event-by-event fluctuation of the initial conditions in hydrodynamics are sizable and should be considered in data analyses. They do depend on the equation of state. Among the quantities examined here, δv_2 is the most sensitive to the equation of state.

In the present work, many important factors have not been considered: baryon-number conservation, strangeness production, resonance decays, continuous emission effects, spectators, etc., which should indeed taken into account in order to get more precise results. Especially, use of the same procedure for the determination of the centrality as used by experimentalists, as in[6], will make the results more directly comparable with data. In any event, we believe that the effects we studied will be present and will be sizable, even with these improvements.

REFERENCES

1. The preliminary version appeared in T. Osada, C.E. Aguiar, Y. Hama and T. Kodama, *Event-by-event analysis of ultra-relativistic heavy-ion collisions in smoothed particle hydrodynamics*, arXiv: nucl-th/0102011.
2. H.J. Drescher, M. Hladik, S. Ostrapchenko, T. Pierog and K. Werner, J.Phys. **G25** (1999) L91; Nucl.Phys. **A661** (1999) 604.
3. C.M. Hung and E.V. Shuryak, Phys. Rev. Lett. **75** (1995) 4003.
4. L.B. Lucy, Ap. J. **82** (1977) 1013; R.A. Gingold and J.J. Monaghan, Mon.Not.R.Astr.Soc. **181** (1977) 375.
5. C.E. Aguiar, T. Kodama, T. Osada and Y. Hama, J.Phys. **G27** (2001) 75, and references therein.
6. STAR Collaboration, K.H. Ackermann *et al.*, Phys.Rev.Lett. **86** (2001) 402.